

Intrinsic spin noise in MgO magnetic tunnel junctions

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We consider two intrinsic sources of noise in ultra-sensitive magnetic field sensors based on MgO magnetic tunnel junctions, coming both from ²⁵Mg nuclear spins (I = 5/2, 10% natural abundance) and S = 1 Mg-vacancies. While nuclear spins induce noise peaked in the MHz frequency range, the vacancies noise peaks in the GHz range. We find that the nuclear noise in submicron devices has a similar magnitude than the 1/f noise, while the vacancy-induced noise dominates in the GHz range. Interestingly, the noise spectrum under a finite magnetic field gradient may provide spatial information about the spins in the MgO layer. © 2013 American Institute of Physics. [http://dx.doi.org/10.1063/1.4791594]

Magnetic tunnel junctions (MTJ) with ferromagnetic electrodes and a MgO tunnel barrier have a very large room temperature tunneling magnetoresistance (TMR).^{1,2} As a result, they are widely used for magnetic sensing applications where room-temperature ultra-high sensitivity, circuit integration, and low fabrication cost are essential. Engineering of multilayer MTJ devices has allowed building devices whose resistance scales linearly with the applied magnetic field. If this linear relation holds at arbitrarily small field, the devices can operate as sensors for magnetic fields as small as permitted by the different sources of noise. In general, noise in MTJ can be classified in two groups, electric and magnetic.^{3–8} The former includes shot-noise, Johnson-Nyquist noise, electric 1/f noise, or noise due to charge trapping in the oxide barrier. The second includes fluctuations in the magnetic orientation of the electrodes due to collective precessional modes, 1/f magnetic noise, domain wall motion, and so on.

More specifically, noise in MgO based MTJ has been widely studied.^{9–13} It has been found that the dominant low frequency 1/f noise scales with the bias voltage, due to resistance fluctuations associated to charge-trapping in the barrier.^{9–11} Furthermore, magnetically dependent 1/f-noise in MgO-MTJ with area larger than a few μ m²,^{9–12} originates from magnetization fluctuations of the free layer (FL).^{9–11}

MgO based TMR sensors with an area of $1\mu m^2$ feature sensitivities of up to pT/\sqrt{Hz} limited by white noise background.^{6,7} This striking sensitivity leads us to address the following intriguing question: to which degree the magnetic field created by spins in the subnanometer thick MgO barrier can be a source of noise that limits the performance of these devices? Or reversing the terms of the question: could the electrical noise of a MgO-MTJ probe the spin noise of the barrier?

The MgO barrier certainly hosts the only stable Mg spinful nuclear isotopes, ²⁵Mg, with nuclear spin, I = 5/2. Thus, taking into account that the lattice constant of the MgO is 4.212 Å, and its natural relative abundance of 10%,¹⁴ the volumetric density of nuclear spins is $\rho_m = 1.32$ spins/nm³. The maximal magnetic field created by one of such nuclear spins, at a distance *l*, reaches $4.3 \text{ nm}^3/l^3 \mu \text{T}$. In addition, the MgO barrier hosts a density of Mg vacancies^{15–18} which have electronic spin *S* = 1, each of which will create a magnetic field 3 orders of magnitude larger.

In standard MTJ sensing devices, one magnetic layer is designed to have its magnetization pinned by exchange coupling to an antiferromagnet while the other is free to rotate, see Fig. 1(a).^{4,6} Then, the relevant figure of merit is given by the sum of all nuclear fields, averaged over the entire FL sensing electrode

$$\vec{\mathcal{B}}(\vec{m}_i) \equiv \frac{1}{V} \int_V \vec{B}[\vec{m}_i](\vec{r}) dV, \qquad (1)$$

where the integral is over the volume V of the detector and $\vec{B}[\vec{m}_i](\vec{r})$ corresponds to the magnetic field created at position \vec{r} by the set of magnetic dipoles $\{\vec{m}_i\}$. If all the nuclear spins were fully polarized, they would create an average field that, for a cylindrical device with R = 100 nm, would lead to $\mathcal{B}_{max} \sim 0.1 \,\mu\text{T}$, which motivates a detailed study of the nuclear spin noise in this system.

At room temperature, the average nuclear spin orientation is vanishingly small, and so it is the average magnetic field they create, but statistical fluctuations of the nuclear spin orientation create magnetic noise. For the calculation of its statistical properties, the following relation between the



FIG. 1. (a) Scheme of a MTJ sensing device. (b) Variation of the standard deviation of the average field in the free layer with the detector radius R for a device with d = 0.5 nm.

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This article is copyrighted as indicated in the abstract. Reuse of AIP content is subject to the terms at: http://scitation.aip.org/termsconditions. Downloaded to IP: 193.145.230.254 On: Wed. 23 Oct 2013 09:58:11 average sensing layer field, Eq. (1), and the nuclear magnetic moments \vec{m}_i is extremely useful:

$$\mathcal{B}_a = \sum_{i,b} \Xi_{ab}(i) m_b(i), \tag{2}$$

where

$$\Xi_{ab}(i) = \frac{\mu_0}{4\pi} \frac{1}{V} \int_V dV \frac{n_b(i)n_a(i) - \delta_{ab}}{\left|\vec{r} - \vec{r}_i\right|^3}.$$
 (3)

 $\Xi_{ab}(i)$ is a geometrical factor that relates the *a* component of the average detector field to the *b* component of the nuclear magnetic moment *i*, with *a*, *b* = *x*, *y*, *z*. The linear relation in Eq. (3) permits relating the quantum statistical properties of the nuclear spins to those of the sensing layer average in a straightforward way, in particular, if one assumes that different nuclear spins are uncorrelated. In this way, the standard deviation of the *a*-magnetic field component created by the fully randomized nuclear spins, defined as $\sigma_{B_a}^2 \equiv (\langle B_a - \langle B_a \rangle)^2$, where the brackets stand for the quantum statistical average, can be written as

$$\sigma_{\mathcal{B}_a}^2 = (g^* \mu_N)^2 I (I+1) \sum_{i,b} \Xi_{ab}(i)^2,$$
(4)

where we have used $\langle m^2 \rangle = g^{*2} \mu_N^2 I(I+1)$, with μ_N the nuclear magneton and g^* the effective g-factor ($g^* \approx 0.342$ for the ²⁵Mg).¹⁹

The quantity $\sigma_{\mathcal{B}_a}$ represents the *a*-component of the nuclear magnetic field noise integrated over the entire frequency range. In addition, if the nuclear spins are randomized, we will find that in cylindrical devices like the one in Fig. 1(a), $\sigma_{\mathcal{B}_x} = \sigma_{\mathcal{B}_y}$. Since we can safely neglect changes in the magnitude of the magnetization, the nuclear noise field can only be efficient in rotating the FL magnetization, which by design of these sensors, can only happen in the plane of the layer. Therefore, only the noise along the in-plane direction x perpendicular to the equilibrium magnetization, will compromise the sensor accuracy. Figure 1(b) shows the numerically calculated $\sigma_{\mathcal{B}_x}$ for two devices with FL thickness w = 2 and 3 nm, and barrier thickness d = 0.5 nm a function of R. Positions \vec{r}_i in the MgO layer have been randomly chosen and we have checked that results do not significantly depend on the random distribution. From Fig. 1, we can extrapolate and get that for $R = 1 \,\mu m$ and $w = 3 \,nm$, $\sigma_{\mathcal{B}_{r}} \approx 10 \,\mathrm{nT}.$

From our numerics, we find that σ_{B_x} grows linearly with 1/R except for very small devices $R \leq 10$ nm. Thus, the relevance of the nuclear spin noise increases for smaller sensors. Notice that from Eq. (4) it is ostensible that $\sigma_{B_a}^2$ scales proportionally to N, the number of nuclear spins in the barrier. This is a consequence of the linear relation in Eq. (2) on one hand, and the linear scaling between the statistical fluctuations of the total magnetic moment and the number of spins.^{20,21} Nevertheless, in our case, the 1/R scaling of the standard deviation of the magnetic field comes from the scaling of integral (3).

In addition to the unavoidable nuclear spin noise, MgO can have a certain density of oxygen and magnesium vacancies.^{15–17,22} The most likely spinfull vacancies in MgO are the Mg vacancies, V_{Mg} , with concentrations that vary between 10^{19} cm⁻³ and 10^{21} cm⁻³.^{15–18} According to density functional calculations,¹⁸ the magnetic moment of these vacancies is $m_{V_{Mg}} \approx 1.9 \,\mu_B$. Whereas the number of vacancies might be smaller than the density of spinfull Mg nuclei, their magnetic moment is also 2000 times larger. Thus, they could also be the source of more spin noise. The analysis of the numerical data shows that, in both cases, $\sigma_{B_a} \propto \sqrt{\langle m^2 \rangle} \sqrt{\rho} / (wR)$ for $R \gg d$, w, so the standard deviation of the field scales with the square root of the barrier spin density, ρ .

We now consider the spectral properties of the nuclear and vacancy magnetic field noise. For that matter, we assume that every nuclear and vacancy spin precesses freely under the influence of the magnetic field created by the ferromagnetic electrodes, \vec{B}_{ext} . Thus, we neglect the mutual coupling between spin centers in the barrier, as well as the nuclearvacancy spin coupling, except for a phenomenological relaxation time T_1 explained below. This is a very good approximation since the magnetic field created by the spin centers is much smaller than the one created by the electrodes. Notice that the precession frequency of nuclear and electronic spins is very different, on account of their different magnetic moments. Then, for a MgO average field of 0.1 T, the nuclear and electronic precession frequencies are in the range of MHz and GHz, respectively.

We assume that the magnetic field felt by the barrier spins is time independent and it only varies in the direction perpendicular to the interfaces (z). This approximation works well as long as the time fluctuations of the magnetic field created by the barrier are slow compared to the barrier spin dynamics. Under these approximations, the correlation function for the detector average at different times, $S_a^2(t)$ $\equiv \langle \mathcal{B}_a(t) \mathcal{B}_a(0) \rangle$, with t > 0, is related to the spin correlation functions as

$$S_{a}^{2}(t) = \sum_{ii',bb'} \Xi_{ab}(i) \Xi_{ab'}(i') \langle m_{b}(i;t)m_{b'}(i';0) \rangle.$$
(5)

The evaluation of this quantity is greatly simplified using the fact that, to a very good approximation, different barrier spins are uncorrelated. Accordingly, the experimentally relevant noise spectrum, $S_x^2(\omega) = \int_{-\infty}^{\infty} e^{-i\omega t} S_x^2(t) dt$, can be expressed as

$$S_x^2(\omega) = \sum_{i,bb'} \Xi_{xb}(i) \Xi_{xb'}(i) \langle m_b(i)m_{b'}(i) \rangle [\omega].$$
(6)

If we quantize the system along the magnetic field orientation at each nuclear spin, and denoting as $|n\rangle$ the nuclear spin eigenstates, the barrier spin spectral function reads, in the limit $k_B T \gg |\vec{m}| B_{ext}$

$$\langle m_b(i)m_{b'}(i)\rangle[\omega] = \frac{\delta_{i,i'}}{(2I+1)} \sum_{nn'} \langle n|m_b|n'\rangle \\ \times \langle n'|m_{b'}|n\rangle \delta(\omega - \omega_{nn'}(i)),$$
(7)

where $\hbar \omega_{nn'}(i) = |\vec{m}| B_{ext}(i)(n - n')$ is the energy of the spin transition $n \to n'$, which depends on local the value of the

external field. Some straightforward algebra permits obtaining the following relation between the spectral noise response $S_x(\omega)$ and σ_{B_x} :

$$\int_{-\infty}^{\infty} S_x^2(\omega) d\omega = \frac{\sigma_{\mathcal{B}_x}^2}{3}.$$
 (8)

As a first approach, let us assume that all the barrier spins feel the same magnetic field intensity. Then, the ²⁵Mg nuclear spins spectral function has a single finite-frequency peak at the Larmor frequency $\omega_B = |\vec{m}| B_{ext}/\hbar$.

Due to its coupling to the environment, the spectral function of a single nuclear spin, Eq. (7), acquires a finite linewidth. We model this by substituting the delta function in Eq. (7) by a Lorentzian function with a width $\delta \omega = 2\pi/T_1$, with T_1 the characteristic relaxation time. Typically, $T_1 \leq 50$ s in bulk MgO at room temperature,²³ and it is expected to be at least 1 ms or larger in surfaces.²⁴ The resulting nuclear noise spectrum is shown in Fig. 2 for two values of *R*. The magnitude of the peak noise associated to the nuclear spins is in the range of nT/Hz^{1/2}, centered in the Larmor frequency (0.5 MHz for $B_{ext} \sim 0.1$ T).

This reported nuclear noise has to be compared with the noise coming from other sources, such as the 1/f noise. We take as a reference a $R = 20 \,\mu\text{m}$ sensor that has a noise level of pT/ $\sqrt{\text{Hz}}$ at 500 kHz.^{25,26} We use the fact that the 1/f noise also scales like 1/R with size, so that, extrapolating down to $R = 100 \,\text{nm}$, the 1/f noise would be 0.4 nT/ $\sqrt{\text{Hz}}$, comparable to the one in Fig. 2(a). Therefore, the contributions of nuclear spin noise and 1/f noise are, under these assumptions, of the same order.

We now consider the noise due to spinful Mg vacancies. If we assume a lower limit for the V_{Mg} concentration of 10^{19} cm⁻³, a small MgO layer of R = 25 nm and d = 0.5 nm will contain more than 10 vacancies. Since the magnetic moment of these vacancies is around $1.9 \mu_B$, at least three orders of magnitude larger than in the ²⁵Mg nuclei, even a single vacancy can produce fluctuations of the magnetic field of the order of μ T for devices with R = 100 nm, see inset of



FIG. 2. (a) Spectral response $S_x(\omega)$ versus frequency $f = \omega/2\pi$ for a detector of radius R = 100 nm (black line) and R = 50 nm (blue line), d = 1 nm, w = 3 nm, $B_{ext} = 0.1$ T, and $T_1 = 10$ ms. (b) Scheme of the variation of the field along a 1 nm thick MgO layer.



FIG. 3. Spectral response $S_x(\omega)$ versus frequency $f = \omega/2\pi$ for a detector of radius R = 100 nm, d = 1 nm, w = 3 nm, and $T_1 = 5 \mu \text{s}$ (black line) and $T_1 = 1 \mu \text{s}$ (red line), containing $320 V_{Mg}$. A magnetic field gradient of 1 mT/nm along the *z*-axis was assumed. Inset shows the integrated standard deviation σ_{B_x} due to a single V_{Mg} located at the center of the MgO layer versus the radius *R*.

Fig. 3. A second consequence of the large difference in magnetic moment with the nuclei is that the corresponding Larmor frequency for typical fields around 0.1T will be in the range of GHz.

The magnitude of the field, which determines the location of the spectral noise peak, is expected to change along the MgO layer since, in general, the magnetization on the FL and pinning layer is different. Magnetic field gradients up to 40 mT/nm have been reported for magnetic disk heads.²⁷ In Fig. 3, we show the effect of a magnetic field gradient of 1 mT/nm. Expectedly, several peaks appear in the spectrum corresponding to different Larmor frequencies, whose position reflects variations of the field across the different Mg atomic planes, see Fig. 2(b).

The different peaks will be resolved if their spectral broadening is smaller than the splitting, $|\vec{m}| \cdot |\partial_z B(z)| d/\hbar \gg 2\pi/T_1$. The relaxation time of these vacancies is much shorter than for the nuclear spins, below 100 μ s.²⁸ Figure 3 shows the spectra corresponding to two different relaxation times, $T_1 = 1$ and 5 μ s. In both cases, the relative height of the different peaks will reflect the abundance of vacancies in each atomic plane of the MgO. Thereby, structural information concerning the distribution of Mg vacancies along the barrier could be inferred from measurements of the noise spectrum.

In conclusion, we have studied the impact of the fluctuating magnetic field created both by the ²⁵Mg nuclear spins and Mg vacancies on a TMR magnetic field sensor with a thin MgO barrier, with circular section of radius *R*. The noise decreases inversely proportional to *R* and it is spectrally peaked at the spin Larmor frequency, determined by the magnetic field in the barriers. Even for a null external field, the magnetic field in the barriers is of the order of 0.1 T. For this value, Larmor frequencies are in the range of 500 kHz for the nuclear spins and 2 GHz for the Mg vacancies. We argue that although the nuclear-induced noise in the 0.5 MHz region is around $1 \text{ nT}/\sqrt{\text{Hz}}$ for devices with R = 100 nm, comparable to the 1/f noise, the vacancies-induced noise should be larger than $1 \text{ nT}/\sqrt{\text{Hz}}$ in the 2 GHz vicinity, well above the 1/f noise. Thus, we expect that Mg vacancies will be a relevant source of magnetic noise, even for very small external field, and thereby potentially relevant for spin sensor applications.

Finally, we show that for a linearly varying magnetic field in the barrier, the noise spectrum can show a series of peaks whose position and height reflects the variations of the magnetic field magnitude and barrier spin density at the different Mg planes. Thus, measurement of this noise, through electrical characterization, could provide some sort of spin imaging of the barrier.

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